TeV-scale Black Holes

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SH, Ben Koch and Marcus Bleicher: hep-ph/0507138, hep-ph/0507140

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Intro

Extra Dimensions Black Holes at the LHC Black Hole Remnants at the LHC

Black Holes as Physics Meeting Point



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Black Hole Remnants at the LHC

Black Holes as Physics Meeting Point

General Relativity

Thermodynamics

Particle Physics

Quantum Field Theory

String Theory

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Why Extra Dimensions

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Why Extra Dimensions

Top down



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Why Extra Dimensions

Top down





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Extra Dimensions



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Top down



Extra Dimensions

"Science may be described as the art of systematic over-simplification."





Karl Popper, The Observer, August 1982

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The ADD-Model

- d+3 space like dimensions (bulk)
- Compactified to radius R
- Only gravitons are allowed into all dimensions
- SM particles bound to 3-dimensional submanifold (brane)



Arkani-Hamed, Dimopoulos and Dvali, Phys. Lett. B **429**, 263 (1998) Antoniadis, Arkani-Hamed, Dimopoulos and Dvali, Phys. Lett. B **436**, 257 (1998) Arkani-Hamed, Dimopoulos and Dvali, Phys. Rev. D **59**, 086004 (1999)

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- + Solves hierarchy problem $m_p^2 = R^d M_f^{d+2}$
- Large radii $1/R \sim {
 m eV}$.. 10 MeV

Black Holes in Extra Dimensions

In large extra dimensions (ADD)

- Gravity stronger at small distances \Rightarrow horizon radius larger
- For mass $M \sim 1 \text{ TeV}$:

 $R_H \sim 2 \times 10^{-38} {
m fm}$ without $R_H \sim 2 \times 10^{-4} {
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Collider produced black holes radius $R_H \ll R$ masses $\sim \text{TeV}$

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 At the LHC partons can come closer than their Schwarzschild horizon — a black hole can be created!



Giddings and Thomas, Phys. Rev. D 65 056010 (2002)

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Cross-section of Black Holes

- Cross section $\sigma \sim \pi R_H^2$ is function of \sqrt{s}
- Threshold $\Theta(M-M_{min})$, one expects $M_{min} \sim M_f$
- Model with colliding wave-packets in Aichelburg-Sexl geometry and examine spacetime for horizons
- Integrate over PDFs for hadron collisions

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Production of Black Holes



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Evaporation of Black Holes

The evaporation proceeds in 3 stages:

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Numerical investigation: black hole event generator CHARYBDIS

Tanaka et al, [arXiv:hep-ph/0411095]; Harris et al, [arXiv:hep-ph/0411022]

Why Black Hole Remanants

Information loss, modified gravity, hair ...

A simple argument:

- Compton-wavelength $\lambda \sim 1/E$, Schwarzschild-radius $R_H \sim E/m_p^2$
- R_H should not be $< \lambda$, because of the uncertainty principle
- ightarrow But would happens for black hole with $E < m_p$

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- \rightarrow But would happens for black hole with $E < m_p$

Also:

• Evaporation of $\lambda > R_H$ not possible



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- Emission of $E \sim 1/\lambda$ larger than $M \sim R_H m_p^2$ is not possible
- Evaporation stops when largest possible wavelength carries already too much energy
 - ightarrow Black hole becomes stable at $M \sim m_p$

Observables of Black Holes

- Multi-jet like events
- Momentum cut-off at $\sim M_f$
- Thermal spectrum \rightarrow yields d and M_f



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Observables of Black Holes

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→ Are significantly modified through the formation of a remnant!

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 p_T -Spectrum of decay products before fragmentation



- Final decay makes an important contribution to p_T-spectrum
- Remnant production makes a significant difference
- After fragmentation, difference is less prominent but still present

$\sum p_T$ -Spectrum of black hole event



• Missing p_T is increased in presence of remnants

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Summary

- Effective models with extra dimensions quantify first effects beyond the standard model
- Black hole production will become important with strong gravitational effects
- Numerical tools for black hole events available
- New: Signatures of black hole event with possible remnant are significantly modified
- \longrightarrow Exciting test for the onset of quantum gravity

More details: hep-ph/0507138, hep-ph/0507140

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